

# Theoretical Micromagnetics

## Lecture Series

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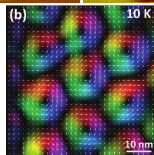
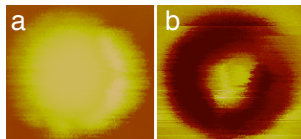
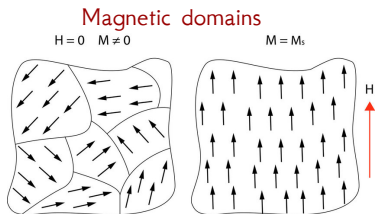
## Lecture 2a. The magnetization vector

Consider a ferromagnet with aligned magnetic moments

The magnetization is the density of magnetic moments  $\mu$  in a volume

$$\mathbf{M} = \frac{\Delta\mu}{\Delta V}, \quad \Delta V \text{ is a small volume.}$$

By applying a strong magnetic field we may align all magnetic moments ("saturate" the magnetization) along the field direction and measure the saturation magnetization  $M_s$ .



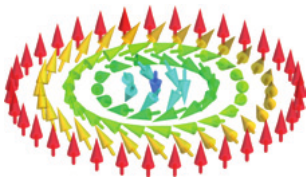
# The Bloch sphere

The magnetization vector takes values on the **Bloch sphere**,  $\mathbf{M}^2 = M_s^2$ .

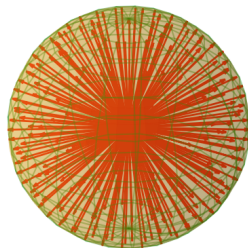
A ferromagnet is described by the magnetization vector  $\mathbf{M} = \mathbf{M}(x, t)$  with (see, Landau, Lifshitz, Pitaevskii, "Statistical Physics II")

$$|\mathbf{M}| = M_s (= \text{const.}).$$

Magnetization configuration



Bloch sphere



# A continuous spin variable

Let us assume a chain of spins which may not be perfectly aligned. The exchange energy depends on the neighbours of each spin  $\mathbf{S}_\alpha$ ,

$$E_{\text{ex}} = -J \sum \mathbf{S}_\alpha \cdot \mathbf{S}_{\alpha+1} = -\frac{J}{2} \sum \mathbf{S}_\alpha \cdot (\mathbf{S}_{\alpha+1} + \mathbf{S}_{\alpha-1}).$$

## A continuum approximation

Consider a small parameter  $\epsilon$  and define ( $\epsilon$  can be defined in different ways)

- A space variable  $x = \epsilon\alpha$  where  $\alpha$  is an integer index ( $\epsilon$  may be the spacing between atoms).
- A continuous field  $\mathbf{S}(x)$  such that  $\mathbf{S}_\alpha = \mathbf{S}(x)$  at the position of each spin  $\alpha$ .

The continuous field  $\mathbf{S}(x)$  is connecting the discrete spins (atoms) of the material.

# Taylor expansion

The advantage of the continuous field is that we can make a

## Taylor approximation

When the distance  $\epsilon$  between spins is small, we have (Taylor expansion)

$$\mathbf{S}_{\alpha\pm 1} \approx \mathbf{S} \pm \epsilon \partial_x \mathbf{S} + \frac{\epsilon^2}{2} \partial_x^2 \mathbf{S}, \quad \mathbf{S}_\alpha \rightarrow \mathbf{S}.$$

This assumes that

- There is a continuous field  $\mathbf{S}(x)$ .
- Neighbouring spins differ only a little.

Example (Use the Taylor approximation in the expression for the exchange energy)

$$E_{\text{ex}} = -\frac{J}{2} \sum \mathbf{S}_\alpha \cdot (\mathbf{S}_{\alpha+1} + \mathbf{S}_{\alpha-1}) \approx \dots$$

# Exchange energy (continuum)

## Exchange energy

Use the Taylor expansion in the exchange energy

$$E_{\text{ex}} = -J \sum \left( |\mathbf{S}|^2 + \frac{\epsilon^2}{2} \mathbf{S} \cdot \partial_x^2 \mathbf{S} \right) \rightarrow -\frac{J}{2} \epsilon \int \mathbf{S} \cdot \partial_x^2 \mathbf{S} dx$$

$$\text{Since } \mathbf{M} \sim \mathbf{S} \text{ we have } E_{\text{ex}} \sim - \int \mathbf{M} \cdot \partial_x^2 \mathbf{M} dx$$

and this gives, by a [partial integration](#)

$$E_{\text{ex}} = \frac{A}{M_s^2} \int \partial_x \mathbf{M} \cdot \partial_x \mathbf{M} dx.$$

- $A$  is the exchange constant (parameter).
- $E_{\text{ex}}$  is non-negative.
- Its minimum (perfect alignment,  $\partial_x \mathbf{M} = 0$ ) lies at zero.
- All directions in space, for  $\mathbf{M}$ , are equivalent.

# The nonlinear $\sigma$ -model

A system with the energy  $E_{\text{ex}}$  and  $\mathbf{M}^2 = \text{const.}$  is called the nonlinear  $\sigma$ -model.

Exercise ( $O(3)$  invariance of  $E_{\text{ex}}$ )

(a) Write  $E_{\text{ex}}$  using the components  $\mathbf{M} = (M_1, M_2, M_3)$ . (b) Consider  $\alpha$  uniform rotation for  $\mathbf{M}$  and show that  $E_{\text{ex}}$  remains invariant.

# Magnetocrystalline anisotropy

Materials are anisotropic in a natural way, e.g., due to the crystal structure. Anisotropic contributions come from relativistic effects. Some types of anisotropy are simply modelled.

## Easy-plane anisotropy

The energy term ( $K > 0$  the anisotropy parameter)

$$E_a = \frac{K}{M_s^2} \int (M_3)^2 dx$$

favours the states where  $\mathbf{M}$  lies on the plane (12), i.e.,  $M_3 = 0$ .

## Easy-axis anisotropy

$$E_a = \frac{K}{M_s^2} \int (M_s^2 - M_3^2) dx$$

favours the states where  $\mathbf{M}$  is fully aligned along the third axis, i.e.,  $M_3 = \pm M_s$  or  $\mathbf{M} = (0, 0, \pm M_s)$ .



## Examples. Minima of anisotropy energy.

### Example (easy-plane anisotropy)

- (a) For the easy-plane anisotropy, give all minimum energy solutions.
- (b) Show that the energy is invariant with respect to rotations of the vector  $\mathbf{M}$  in the (12) plane.

### Example (easy-axis anisotropy)

- (a) Write the easy-axis anisotropy formula in a manifestly non-negative form to show that  $E_a \geq 0$ .
- (b) Give all minimum energy solutions (based on that formula).

# Energy and length scales

In three-dimensions (3D), we have the exchange energy

$$E_{\text{ex}} = \frac{A}{M_s^2} \int \partial_\mu \mathbf{M} \cdot \partial_\mu \mathbf{M} d^3x, \quad \mu = 1, 2, 3.$$

Question (Write explicitly the exchange energy density)

*Note that summation is implied for the repeated index  $\mu$ .*

Total energy

In a simple model we assume a ferromagnet with exchange and anisotropy energy. For a 3D magnet,

$$E = E_{\text{ex}} + E_{\text{a}} = \frac{A}{M_s^2} \int \partial_\mu \mathbf{M} \cdot \partial_\mu \mathbf{M} d^3x + \frac{K}{M_s^2} \int (M_s^2 - M_3^2) d^3x.$$

Units for the physical constants  $A : \text{J/m}$ ,  $K : \text{J/m}^3$ .

## Dimensional analysis

The length scale of this model

The two energy terms indicate a natural length scale

$$\ell_{\text{DW}} = \sqrt{\frac{A}{K}}.$$

Example (For  $A = 10^{-11}$  J/m,  $M_s = 10^6$  A/m,  $K = 4 \times 10^5$  J/m<sup>3</sup>)

We calculate  $\ell_{\text{DW}} = 5 \times 10^{-9}$  m = 5 nm.

We define the dimensionless magnetization according to

$$\mathbf{m} = \frac{\mathbf{M}}{M_s}, \quad \mathbf{m}^2 = 1$$

and we have the energy

$$\begin{aligned} E &= A \int \frac{\partial \mathbf{m}}{\partial x_\mu} \cdot \frac{\partial \mathbf{m}}{\partial x_\mu} d^3x + K \int (1 - m_3^2) d^3x \\ &= K \left[ \ell_{\text{DW}}^2 \int \frac{\partial \mathbf{m}}{\partial x_\mu} \cdot \frac{\partial \mathbf{m}}{\partial x_\mu} d^3x + \int (1 - m_3^2) d^3x \right] \end{aligned}$$

## Dimensionless form of energy

We define dimensionless space variables (i.e., scale space by  $\ell_{\text{DW}}$ )

$$x_\mu = \xi_\mu \ell_{\text{DW}}$$

and have the energy

$$E = (K\ell_{\text{DW}}^3) \left[ \int \partial_\mu \mathbf{m} \cdot \partial_\mu \mathbf{m} d^3\xi + \int (1 - m_3^2) d^3\xi \right].$$

We write  $K\ell_{\text{DW}}^3 = A\ell_{\text{DW}}$ , and re-instate the usual variable  $\xi \rightarrow x$  to get

$$E = (2A\ell_{\text{DW}}) \left[ \frac{1}{2} \int \partial_\mu \mathbf{m} \cdot \partial_\mu \mathbf{m} d^3x + \frac{1}{2} \int (1 - m_3^2) d^3x \right].$$

The natural energy scale is  $(2A\ell_{\text{DW}})$

### Remark

This scaled energy form has no free parameter.

## Lecture 2b. Derivation of the time-independent equation

### Static configurations of the magnetization

The magnetization  $\mathbf{m}(\mathbf{x})$  of the material reduces to a configuration that minimizes the magnetic energy  $E(\mathbf{m})$ .

The equation for  $\mathbf{m}(\mathbf{x})$  is obtained as the Euler-Lagrange equation for the minimization of the energy, with the constraint

$$\mathbf{m}^2(\mathbf{x}) = 1.$$

The constraint is imposed via a *Lagrange multiplier*  $\lambda(\mathbf{x})$ . [Raj, Sec. 3.3][FG, Sec. 12.2]

For a demonstration, we consider the exchange interaction and we have to extremize the functional

$$L[\mathbf{m}] = \int d^3x \underbrace{\left[ \frac{1}{2} \partial_\mu \mathbf{m} \cdot \partial_\mu \mathbf{m} + \frac{\lambda(\mathbf{x})}{2} (1 - \mathbf{m}^2) \right]}_{\mathcal{L}}.$$

# The Euler-Lagrange equation

The functional  $L$  is minimized for  $\mathbf{m}(\mathbf{x})$  that satisfies the Euler-Lagrange equation

$$-\frac{\delta L}{\delta \mathbf{m}} = 0 \Rightarrow \frac{d}{dx_\mu} \left( \frac{\partial \mathcal{L}}{\partial_\mu \mathbf{m}} \right) - \frac{\partial \mathcal{L}}{\partial \mathbf{m}} = 0.$$

We calculate

$$-\frac{\delta L}{\delta \mathbf{m}} = \frac{d}{dx_\mu} (\partial_\mu \mathbf{m}) + \lambda \mathbf{m} = \partial_\mu \partial_\mu \mathbf{m} + \lambda \mathbf{m} = 0$$

or

$$\Delta \mathbf{m} + \lambda \mathbf{m} = 0.$$

We multiply the above by  $\mathbf{m}$  in order to obtain the Lagrange multiplier,

$$\mathbf{m} \cdot \Delta \mathbf{m} + \lambda \mathbf{m} \cdot \mathbf{m} = 0 \Rightarrow \lambda = -\mathbf{m} \cdot \Delta \mathbf{m}$$

and we use this to eliminate  $\lambda$  in the field equation

$$\Delta \mathbf{m} - (\mathbf{m} \cdot \Delta \mathbf{m})\mathbf{m} = 0 \Rightarrow \mathbf{m} \times (\mathbf{m} \times \Delta \mathbf{m}) = 0.$$

The latter is equivalent to

$$\mathbf{m} \times \Delta \mathbf{m} = 0.$$

### Quiz

Equation for the minimization of the exchange energy. Give an example of solution for the 1D equation

$$\mathbf{m} \times \partial_x^2 \mathbf{m} = 0.$$

# The Landau-Lifshitz (LL) equation - static sector

## Form of the Landau-Lifshitz equation

Let us assume an energy functional  $E(\mathbf{m})$ . We find

$$\mathbf{m} \times \mathbf{f} = 0, \quad \mathbf{f} = -\frac{\delta E}{\delta \mathbf{m}}.$$

- For  $\mathbf{f} = \mathbf{h}$  we recover the standard equation of magnetism for a magnetic moment  $\mathbf{m}$  in an external magnetic field  $bh$ .
- For  $E = E_{\text{ex}} = \frac{1}{2} \int \partial_x \mathbf{m} \cdot \partial_x \mathbf{m} dx$  we have  $\mathbf{f} = -\frac{\delta E}{\delta \mathbf{m}} = \partial_x^2 \mathbf{m}$ .
- Solutions are  $\mathbf{m}$  such that  $\mathbf{m} \parallel \mathbf{f}$ .

## Exercise (Static Landau-Lifshitz equation)

Assume an energy functional  $E$  and derive the static Landau-Lifshitz equation under the constraint  $\mathbf{m}^2 = 1$ .



# The LL equation - exchange and uniaxial anisotropy

Energy (exchange and easy-axis anisotropy)

$$E = \int \epsilon dx = \frac{1}{2} \int \partial_x \mathbf{m} \cdot \partial_x \mathbf{m} dx + \frac{1}{2} \int (1 - m_3^2) dx.$$

The variational derivative

$$\mathbf{f} = -\frac{\delta E}{\delta \mathbf{m}} = \frac{d}{dx} \left( \frac{\partial \epsilon}{\partial (\partial_x \mathbf{m})} \right) - \frac{\partial \epsilon}{\partial \mathbf{m}} = \partial_x^2 \mathbf{m} + m_3 \hat{\mathbf{e}}_3.$$

Landau-Lifshitz equation for exchange and easy-axis anisotropy

$$\mathbf{m} \times \underbrace{(\Delta \mathbf{m} + m_3 \hat{\mathbf{e}}_3)}_{\mathbf{f}} = 0.$$

Quiz

on the model with exchange and anisotropy.

### Question (Find the uniform solutions)

- Uniform solutions are those for space-independent  $\mathbf{m}$ .

In this case,  $\mathbf{f} = m_3 \hat{\mathbf{e}}_3$ . In order to have a solution, we need  $\mathbf{m} \parallel \mathbf{f}$ , that is,  $\mathbf{m} \parallel \hat{\mathbf{e}}_3 \Rightarrow \mathbf{m} = \pm \hat{\mathbf{e}}_3$  (pointing in the north or south pole).

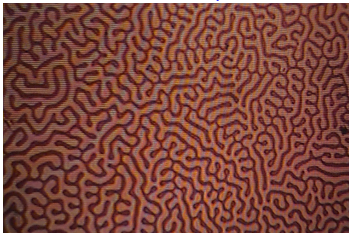
- The uniform solution is called the *ferromagnetic state*.

### Question (easy-plane anisotropy)

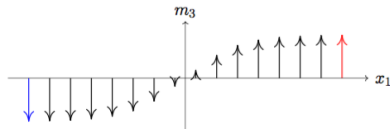
What is the static Landau-Lifshitz equation for easy-plane anisotropy?

## Lecture 2c. The magnetic domain wall (DW)

Domain pattern



Sketch of domain wall



### A transition layer

- Magnetic domains are regions where the magnetization is almost uniform.
- A domain wall is the magnetization configuration between two uniform states with different magnetization.

# The spherical parametrization for the magnetization

## The spherical angles $\Theta, \Phi$

We can explicitly resolve the constraint  $\mathbf{m}^2 = 1$ ,

$$m_1 = \sin \Theta \cos \Phi, \quad m_2 = \sin \Theta \sin \Phi, \quad m_3 = \cos \Theta.$$

In a model with easy-axis anisotropy, we have two ground states,  $\mathbf{m} = (0, 0, \pm 1)$ , or  $\Theta = 0, \pi$  (north and south pole of the sphere).

We look for a topological soliton connecting the north and the south pole

We confine ourselves to the one-dimensional case  $\mathbf{m} = \mathbf{m}(x)$ .

We try the simplest possibility of a meridian on the Bloch sphere

$$\Theta = \Theta(x), \quad \Phi = \phi_0 : \text{const.}$$

# A magnetic domain wall on the Bloch sphere

## Example (Bloch wall)

For  $\phi_0 = \pi/2$  we have

$$m_1 = 0, \quad m_2(x) = \sin \Theta(x), \quad m_3(x) = \cos \Theta(x).$$

- Draw a DW on the Bloch sphere.
- Consider the variation of the vector  $\mathbf{m}$  in the space variable  $x$ .

## A magnetic domain wall (DW)

The Landau-Lifshitz equation for  $\mathbf{m}(x)$  (for exchange and easy-axis anisotropy with anisotropy parameter  $k^2$ )

$$\mathbf{m} \times (\mathbf{m}'' + k^2 m_3 \hat{\mathbf{e}}_3) = 0 \Rightarrow \begin{cases} m_2 m_3'' - m_3 m_2'' + k^2 m_2 m_3 = 0 \\ m_3 m_1'' - m_1 m_3'' - k^2 m_1 m_3 = 0 \\ m_1 m_2'' - m_2 m_1'' = 0 \end{cases}$$

Choose the case

$$m_1 = 0, \quad m_2 = \sin \Theta, \quad m_3 = \cos \Theta.$$

thus

$$m_2'' = -\sin \Theta \Theta'^2 + \cos \Theta \Theta'', \quad m_3'' = -\cos \Theta \Theta'^2 - \sin \Theta \Theta''.$$

The first equation gives (the other two are trivially satisfied)

$$\Theta'' - k^2 \sin \Theta \cos \Theta = 0.$$

Multiply by  $2\Theta'$

$$[(\Theta')^2 - k^2 \sin^2 \Theta]' = 0 \Rightarrow (\Theta')^2 - k^2 \sin^2 \Theta = C.$$

There are many solutions for the one-dimensional equation

We are only interested in *localized solutions*. We consider uniform domains for  $|x| > 0$ , therefore, we require  $\Theta = 0, \pi$  and  $\Theta' = 0$  at  $x = \pm\infty$ .

From the condition at  $x \rightarrow \pm\infty$  we get  $C = 0$  and we have

$$\Theta' = \pm k \sin \Theta.$$

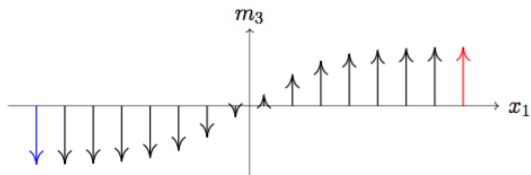
The solution of the latter is

$$e^{\pm kx} = \pm \tan \left( \frac{\Theta}{2} \right).$$

Check that (for the plus signs)

- For  $x \rightarrow -\infty$  we have  $\Theta = 0$  (north pole).
- For  $x \rightarrow \infty$  we have  $\Theta = \pi$  (south pole).

## Domain wall details



### Static domain wall (DW)

Use trigonometric identities (for the half angle)

$$m_1 = \frac{1}{\cosh(kx)} \cos \phi_0, \quad m_2 = \frac{1}{\cosh(kx)} \sin \phi_0, \quad m_3 = \tanh(kx).$$

That is valid for boundary conditions  $\mathbf{m}(x = \pm\infty) = (0, 0, \pm 1)$ .

The figure shows  $m_3(x)$  for a domain wall with  $\phi_0 = \pm\pi/2$ .

The width of the domain wall can be considered to be  $1/k$ , i.e.,

$$\ell_{\text{DW}} = \sqrt{A/K}.$$



# Bloch and Néel walls

There are many domain wall solutions

We get a different domain wall solution for every  $0 \leq \phi_0 < 2\pi$ . Within this model, the energy is the same for all walls.

Bloch wall, choose  $\phi_0 = \pm\pi/2$

$$m_1 = 0, \quad m_2 = \pm \frac{1}{\cosh(kx)}, \quad m_3 = \tanh(kx).$$

Néel wall, choose  $\phi_0 = 0, \pi$

$$m_1 = \pm \frac{1}{\cosh(kx)}, \quad m_2 = 0, \quad m_3 = \tanh(kx).$$

## Lecture 2e. Magnetostatic field

### Maxwell's equations

A ferromagnet produces a magnetic field  $\mathbf{H}_m$ . For static configurations  $\mathbf{M}$ , this is given by Maxwell's equations omitting time derivatives

$$\nabla \cdot \mathbf{B} = 0, \quad \nabla \times \mathbf{H}_m = 0, \quad \mathbf{B} = \mu_0(\mathbf{H}_m + \mathbf{M}).$$

Apply the normalization

$$\mathbf{h}_m = \frac{\mathbf{H}_m}{M_s}$$

and write

$$\nabla \cdot (\mathbf{h}_m + \mathbf{m}) = 0, \quad \nabla \times \mathbf{h}_m = 0.$$

This is called the **magnetostatic field**  $\mathbf{h}_m$ , because time derivatives have been neglected in Maxwell's equations.

# Source of a magnetostatic field

## Magnetic field due to $\mathbf{m}$

Write Maxwell's equations as

$$\nabla \cdot \mathbf{h}_m = -\nabla \cdot \mathbf{m}, \quad \nabla \times \mathbf{h}_m = 0.$$

Thus, the magnetic field source is  $-\nabla \cdot \mathbf{m}$ .

Note the similarity between the equations for the magnetostatic field

with those for the field  $\mathbf{E}$  of a charge density  $\rho$  in electrostatics. They are identical under the correspondence

- $\rho \rightarrow -\nabla \cdot \mathbf{m}$ .
- $\mathbf{E} \rightarrow \mathbf{h}_m$ .

## Quiz. Examples in simple geometries.

### Example (Magnetic field of an infinite cylinder)

Consider an infinite cylinder that is uniformly magnetized along its axis ( $\mathbf{m} = \hat{\mathbf{e}}_3$ ). What is the magnetic field produced?

### Example (Magnetic field in a thin film)

Consider a thin film uniformly magnetized perpendicular to the film plane ( $\mathbf{m} = \hat{\mathbf{e}}_3$ ). What is the magnetic field produced?

# Magnetostatic field: examples

Example (Infinitely elongated cylinder uniformly magnetized along its axis)

The magnetostatic field is (note that  $\nabla \cdot \mathbf{m} = 0$ )

$$\mathbf{h}_m = 0.$$

Example (Thin film uniformly magnetized)

Consider an infinite thin film in the  $xy$  plane uniformly magnetized perpendicular to the plane,  $\mathbf{m} = \hat{\mathbf{e}}_z$ . The magnetostatic field is

$$\mathbf{h}_m = -\mathbf{m} = -\hat{\mathbf{e}}_z.$$

Solutions of the latter type appear in examples in textbooks, e.g., in the case of the field in an ideal capacitor.

## Quiz. Magnetostatic field of a wall.

Bloch wall, choose  $\phi_0 = \pm\pi/2$

$$m_1 = 0, \quad m_2 = \pm \frac{1}{\cosh(kx)}, \quad m_3 = \pm \tanh(kx).$$

This gives  $\nabla \cdot \mathbf{m} = 0$  and thus produces no magnetic field. It minimizes the magnetostatic energy (not included in our model so far).

Néel wall, choose  $\phi_0 = 0, \pi$

$$m_1 = \pm \frac{1}{\cosh(kx)}, \quad m_2 = 0, \quad m_3 = \pm \tanh(kx).$$

This gives  $\nabla \cdot \mathbf{m} = m'_1 \neq 0$  and thus magnetic field is produced. This is added to the domain wall energy.

# A nontrivial example with a simple solution

## Example

In the case of  $\mathbf{m} = \mathbf{m}(x)$ , depending only on one space variable, we have the solution

$$\mathbf{h}_m(x) = -m_x(x)\hat{\mathbf{e}}_x.$$

This is because

- The equation  $\nabla \cdot \mathbf{h}_m = -\nabla \cdot \mathbf{m}$  reduces to the 1D form  $\partial_x h_x = -\partial_x m_x$  and it is satisfied.
- We assume  $m_x(x) = 0$  at  $x \rightarrow \pm\infty$ , thus  $\mathbf{h}_m$  satisfies the boundary condition  $\mathbf{h}_m(\pm\infty) = 0$ .
- At  $y, z \rightarrow \pm\infty$  we do not impose a particular boundary condition (we only assume that  $\mathbf{h}_m$  does not depend on  $y, z$ ).

# The exchange length

## The magnetostatic energy

$$E_m = \frac{1}{2} \mu_0 \int \mathbf{M} \cdot \mathbf{H}_m d^3x.$$

A typical energy density is  $\frac{1}{2} \mu_0 M_s^2$  (in J/m<sup>3</sup>).

Comparison of exchange and magnetostatic energy gives rise to the definition of the **exchange length**

$$\ell_{\text{ex}} = \sqrt{\frac{2A}{\mu_0 M_s^2}}.$$

## Example (Exchange length for Permalloy)

For Permalloy,  $A = 1.3 \times 10^{-11}$  J/m,  $M_s = 0.69 \times 10^6$  A/m. We find  $\ell_{\text{ex}} = 6.59$  nm.



## Rationalise using the exchange length

We may define dimensionless variables according to

$$x \rightarrow x \ell_{\text{ex}}, \quad \mathbf{h}_m = \frac{\mathbf{H}_m}{M_s}.$$

We get the energy, in units of  $2A\ell_{\text{ex}}$ ,

$$E = \frac{1}{2} \int \partial_\mu \mathbf{m} \cdot \partial_\mu \mathbf{m} d^3x + \frac{k^2}{2} \int (1 - m_3^2) d^3x + \frac{1}{2} \int \mathbf{m} \cdot \mathbf{h}_m d^3x$$

where we defined (the "quality factor")

$$k^2 = \frac{2K}{\mu_0 M_s^2}.$$

### Remark

This form of the energy has only one parameter  $k^2$ , the scaled (dimensionless) anisotropy.